

# AdS/CFT, hydrodynamics and beyond

Romuald A. Janik

Jagiellonian University  
Kraków

M. Heller, RJ, P. Witaszczyk, M. Spaliński, work in progress

*Disclaimer:*

I will not talk directly about QCD...

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*... The N=4 super Yang-Mills plasma is finally concluded to be very similar the QCD plasma of gluons and light quarks. The differences mostly reflect different numbers of degrees of freedom in the two systems.*

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## Outline

**Key question**

$\mathcal{N} = 4$  plasma versus QCD plasma

**The AdS/CFT setup**

**Small excitations of a uniform static (strongly coupled) plasma**

**Nonlinear regime**

**Beyond hydrodynamics – towards a 4D description**

## Key question:

Understand the **real-time** properties of strongly coupled gauge theory plasma

Consider plasma in  $\mathcal{N} = 4$  Super-Yang-Mills theory  
Use AdS/CFT methods...

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## $\mathcal{N} = 4$ plasma versus QCD plasma

### Similarities:

- ▶ Deconfined phase
- ▶ Strongly coupled
- ▶ **No supersymmetry!**

### Differences:

- ▶ No running coupling  $\rightarrow$  Even at very high energy densities the coupling remains strong
- ▶ (Exactly) conformal equation of state  $\rightarrow$  Differences close to  $T_c$ , no bulk viscosity... (but not that different around  $1.5 - 2.5 T_c$ )
- ▶ No confinement/deconfinement phase transition  $\rightarrow$  Plasma fireball cools indefinitely

Some of these differences can be lifted in more complicated versions of the AdS/CFT correspondence

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**How to describe strongly coupled plasma using AdS/CFT?**

## The AdS/CFT setup ( $T_{\mu\nu}$ – geometry dictionary)

**Method:** Describe the (possibly time dependent) strongly coupled plasma system through a dual 5D geometry — given e.g. by

$$ds^2 = \frac{g_{\mu\nu}(x^\rho, z) dx^\mu dx^\nu + dz^2}{z^2} \equiv g_{\alpha\beta}^{5D} dx^\alpha dx^\beta$$

i) use Einstein's equations as field equations for  $g_{\mu\nu}(x^\rho, z)$

$$R_{\alpha\beta} - \frac{1}{2} g_{\alpha\beta}^{5D} R - 6 g_{\alpha\beta}^{5D} = 0$$

ii) read off  $\langle T_{\mu\nu}(x^\rho) \rangle$  from the numerical metric  $g_{\mu\nu}(x^\rho, z)$

$$g_{\mu\nu}(x^\rho, z) = \eta_{\mu\nu} + z^4 g_{\mu\nu}^{(4)}(x^\rho) + \dots \quad \langle T_{\mu\nu}(x^\rho) \rangle = \frac{N_c^2}{2\pi^2} \cdot g_{\mu\nu}^{(4)}(x^\rho)$$



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## Example: Static uniform plasma

- ▶ Start from a constant diagonal energy momentum tensor (with  $E = 3p$ )

$$T_{\mu\nu} = \begin{pmatrix} E & 0 & 0 & 0 \\ 0 & p & 0 & 0 \\ 0 & 0 & p & 0 \\ 0 & 0 & 0 & p \end{pmatrix}$$

- ▶ Solve Einstein's equations with the above boundary condition for  $g_{\mu\nu}(x^\rho, z)$ ...
- ▶ The result is a **black hole** geometry

$$ds^2 = -\frac{(1 - z^4/z_0^4)^2}{(1 + z^4/z_0^4)z^2} dt^2 + (1 + z^4/z_0^4) \frac{dx^2}{z^2} + \frac{dz^2}{z^2}$$

with  $z_0$  expressed in terms of  $E$   $(E = \frac{3N_c^2}{2\pi^2 z_0^4})$

- ▶ Hawking temperature  $T_H = \frac{\sqrt{2}}{\pi z_0} \equiv$  gauge theory temperature

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**What kind of excitations propagate in a static uniform strongly coupled plasma system?**

## Answer 1: Hydrodynamics

- ▶ The energy-momentum tensor  $T_{\mu\nu}$  is expressed in terms of a local temperature  $T$  and flow velocity  $u^\mu$
- ▶  $T_{\mu\nu}$  is expressed as an expansion in the gradients of the flow velocities (shown here for  $\mathcal{N} = 4$  SYM)

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^\mu u^\nu)}_{\text{perfect fluid}} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{\text{viscosity}} + \underbrace{(\pi T^2) \left( \log 2 T_{2a}^{\mu\nu} + 2 T_{2b}^{\mu\nu} + (2 - \log 2) \left( \frac{1}{3} T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu} \right) \right)}_{\text{second order hydrodynamics}}$$

- ▶ Consider small perturbations

$$T_{\mu\nu} = \begin{pmatrix} E & 0 & 0 & 0 \\ 0 & p & 0 & 0 \\ 0 & 0 & p & 0 \\ 0 & 0 & 0 & p \end{pmatrix} + \delta T_{\mu\nu} e^{-i\omega t + ikx}$$

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## Answer 1: Hydrodynamics

- ▶ If  $T_{\mu\nu}$  is described by (1<sup>st</sup> order viscous) hydrodynamics then one can derive dispersion relation of long wavelength modes from hydrodynamic equations:

shear modes:

$$\omega_{shear} = -i \frac{\eta}{E + p} k^2$$

sound modes:

$$\omega_{sound} = \frac{1}{\sqrt{3}} k - i \frac{2}{3} \frac{\eta}{E + p} k^2$$

- ▶ If we were to include terms in  $T_{\mu\nu}$  with more derivatives (higher order viscous hydrodynamics), we would get terms with higher powers of  $k$  in the dispersion relations...
- ▶ Hypothetical resummed *all-order* hydrodynamics would predict the full dispersion relation for these modes  $\omega_{shear}(k)$ ,  $\omega_{sound}(k)$

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- ▶ Small disturbances of the uniform static plasma  $\equiv$  small perturbations of the black hole metric ( $\equiv$  quasinormal modes (QNM))

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- ▶ Dispersion relation fixed by linearized Einstein's equations. Results for the sound channel

from Kovtun, Starinets hep-th/0506184

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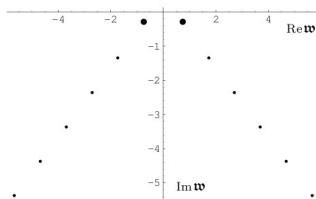


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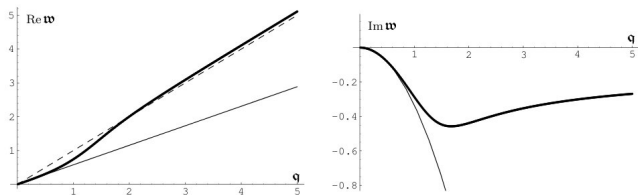
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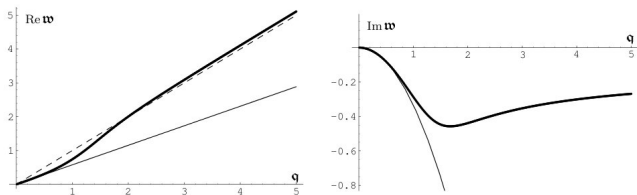
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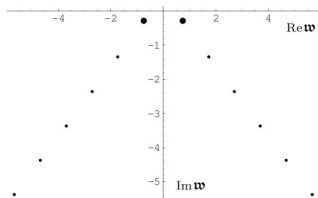
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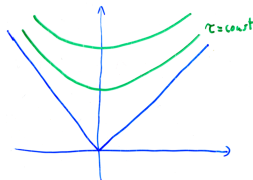
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Bjorken '83

Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.



- ▶ In a conformal theory,  $T_{\mu}^{\mu} = 0$  and  $\partial_{\mu} T^{\mu\nu} = 0$  determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function  $\varepsilon(\tau)$ , the energy density at mid-rapidity.
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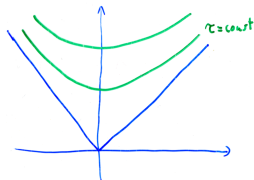
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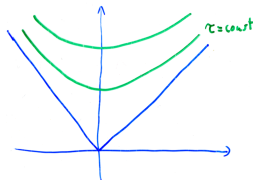
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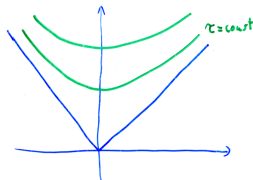
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$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

- ▶ Previously, we normalized our initial data by setting

$$T_{eff}(\tau = 0) = 1$$

but this is generically unknown in realistic heavy-ion collisions...

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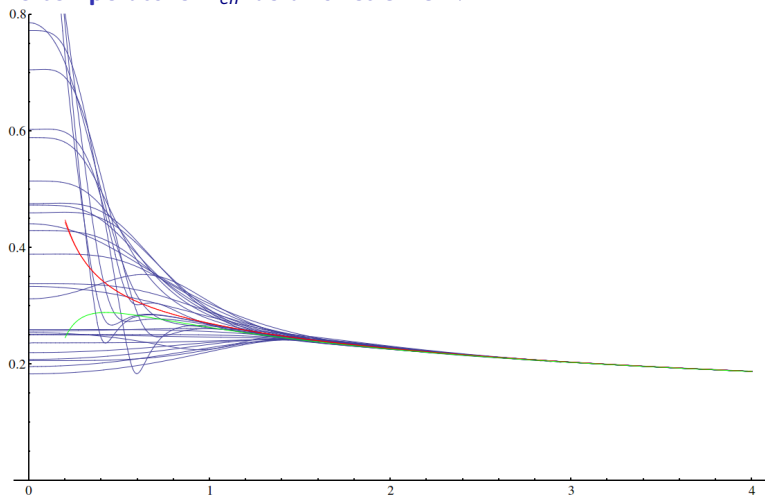
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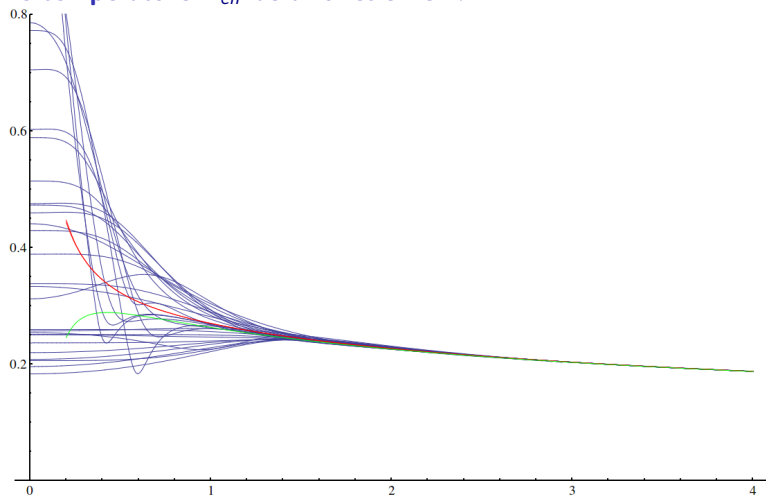


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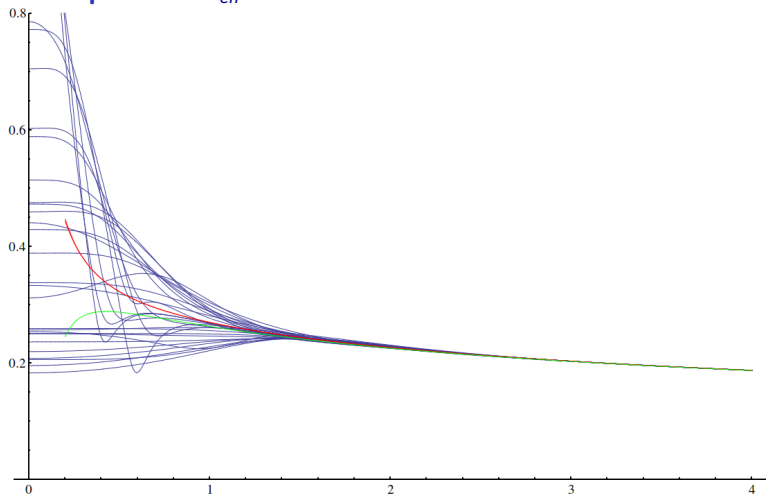


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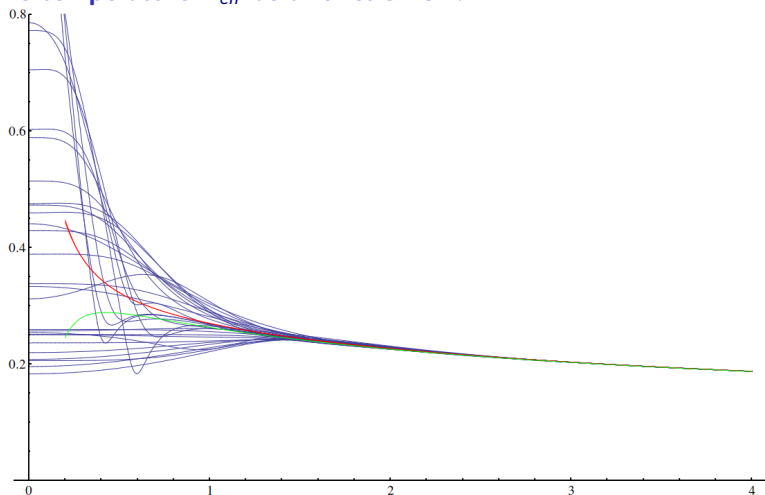


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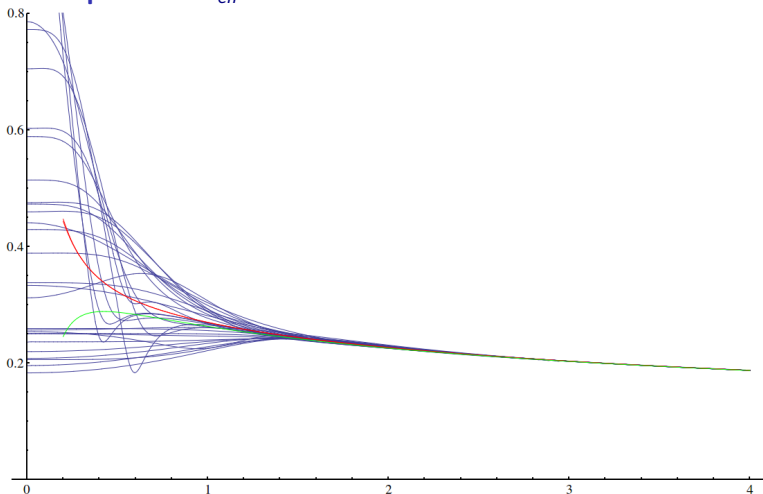


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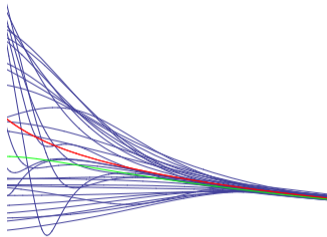
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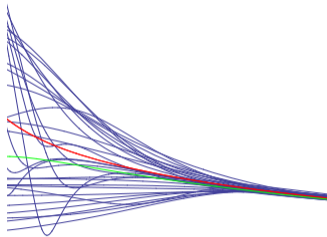
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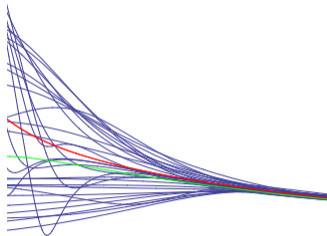
**How to model deviations from (all-order) hydrodynamics?**

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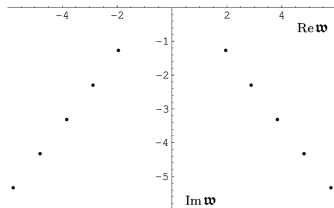
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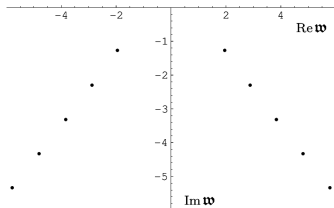
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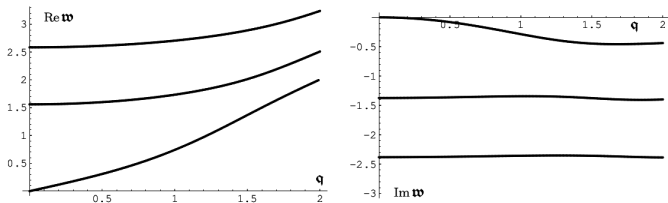
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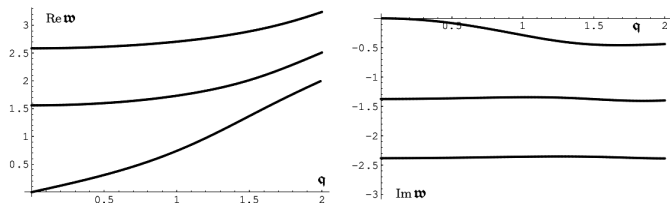
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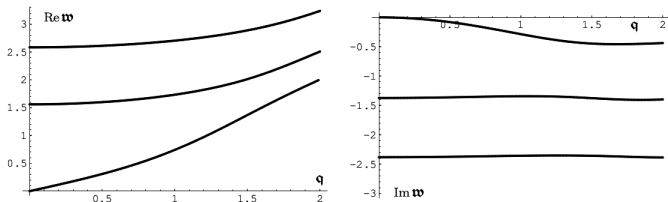
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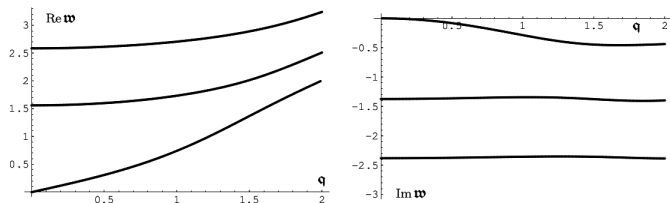
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- ▶ The equation in a generic hydrodynamic background should *locally* look like the equation above...
- ▶ We should 'covariantize' the time derivative...  
Simplest choice:

$$\frac{1}{T} \frac{d}{dt} \longrightarrow \mathcal{D} \equiv \frac{1}{T} u^\mu \partial_\mu$$

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## The case of $T_{\mu\nu}$ (nonequilibrium extension of hydrodynamics)

- ▶ We aim at including the lowest nonhydrodynamic mode

**static uniform plasma background**

- ▶ We have the following tensor structures relative to the rest frame of the fluid:

$$T^{tt} \quad T^{ti} \quad T^{ij}$$

- ▶ Energy-momentum conservation  $\partial_\mu T^{\mu\nu} = 0$  relates the time evolution of the first two to other components...
- ▶ It is enough to just consider a transverse, traceless tensor  $\Pi_{\mu\nu}$
- ▶ This is similar to Israel-Stewart but the equations of motion are different — they have to be second order
- ▶ Again the equations are parametrized by the real and imaginary part of the QNM frequencies:  $\omega_R$  and  $\omega_I$

Proceed to general hydrodynamic flows...

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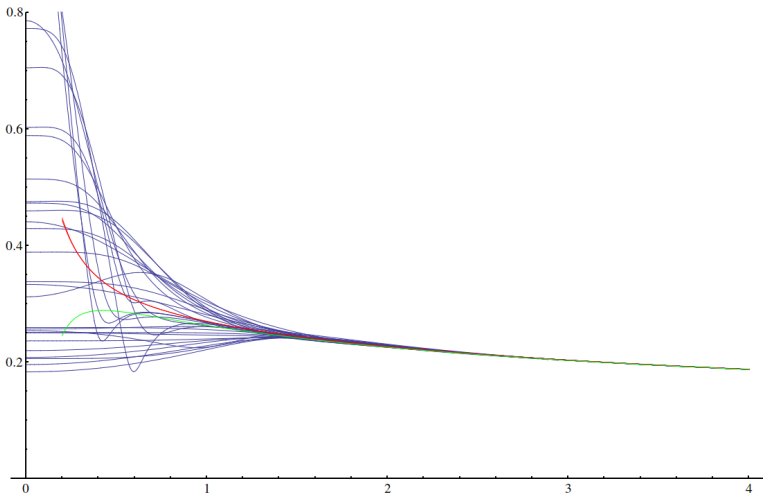
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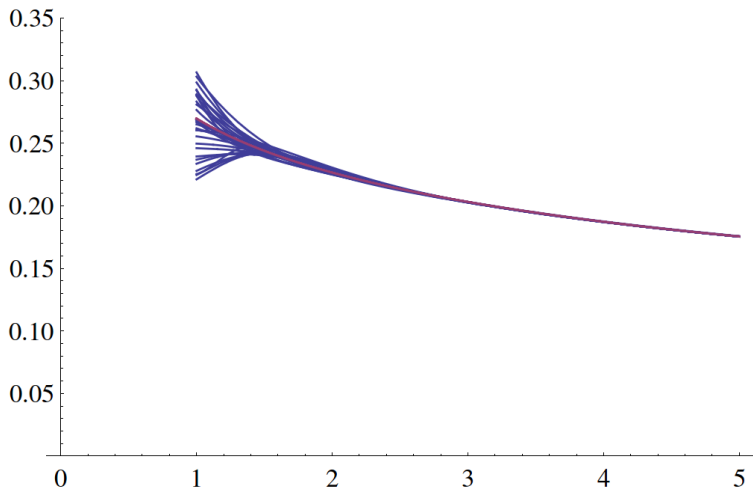
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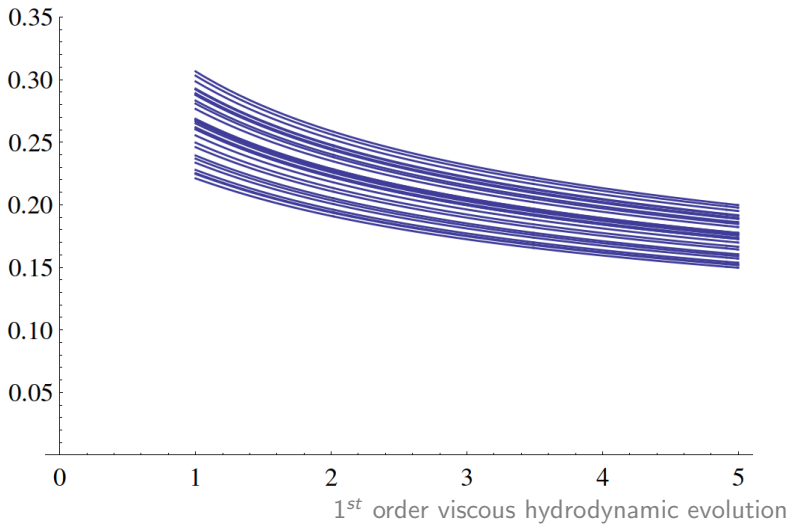
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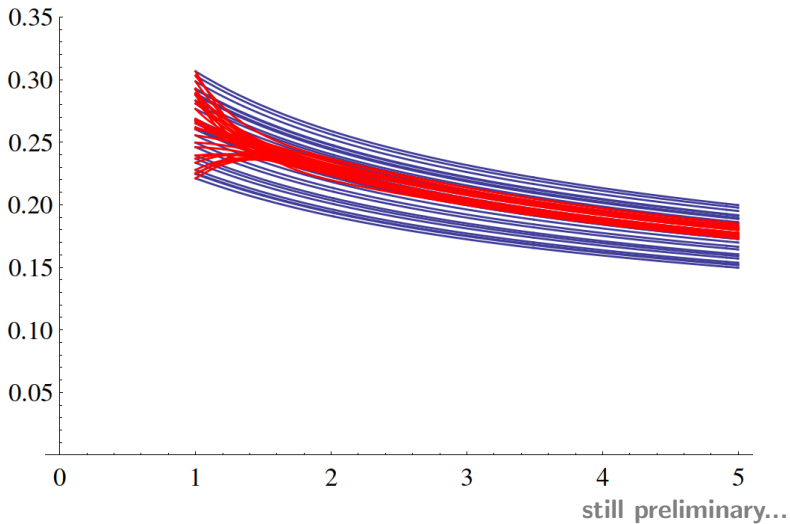
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## Conclusions

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- ▶ Far from equilibrium these nonhydrodynamic degrees of freedom are very important for the dynamics
- ▶ We aim to incorporate the dynamics of lowest nonhydrodynamic degrees of freedom into a purely 4D description
- ▶ Key role of quasinormal frequencies...
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- ▶ **Question:** What are their values in strongly coupled QCD quark-gluon plasma??

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